

# Results from the search for neutrinoless double-beta decay of Te-130 with CUORE-0

Kyungeun E. Lim (on behalf of the CUORE collaboration)



Apr. 14, 2015, LNS Seminar, MIT



#### What we don't know about Neutrinos Neutrino Mass Splitting $m^2$ $m^2$ v<sub>e</sub> 100 - 500 meV $\nu_{\mu}$ $v_{\tau}$ Degenerate $m_3^2$ $m_2^2$ $\int solar \sim 7 \times 10^{-5} eV^2$ $m_1^{2}$ atmospheric $\sim 2 \times 10^{-3} eV^2$ atmospheric $\sim 2 \times 10^{-3} eV^2$ $m_2^2$ solar~7×10<sup>-5</sup>eV<sup>2</sup> $m_1^2$ $m_{3}^{2}$ ? 9 0 0 Normal Hierarchy **Inverted Hierarchy** Is the neutrino its own antiparticle? Rep. Prog. Phys. 76, 056201 (2013)

#### Outline



- Neutrinoless double-beta decay ( $0\nu\beta\beta$ ) search
- CUORE : An array of TeO<sub>2</sub> bolometers
- **Our Comparison of the search of the search** 
  - CUORE-0 : Detector
  - CUORE-0 : Performance and Background
  - CUORE-0 : Results

## Summary

#### Neutrino(less) double-beta decay

0νββ









- Allowed in SM
- Observed in several nuclei  $(T_{1/2}^{2v} \sim 10^{18} - 10^{21} \text{ yr})$
- Beyond SM
- Hypothetical process only if  $v = \overline{v}$  and  $m_v > 0$
- Observation of  $0\nu\beta\beta$
- will establish that neutrinos are Majorana Particles ( $v = \overline{v}$ )
- 2. demonstrate lepton number is not a symmetry of nature
- 3. will provide indirect info about the v mass
- 4. may provide info about the mass hierarchy in combination with direct neutrino mass measurement



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#### Observation of $0\nu\beta\beta$

- I. will establish that neutrinos are Majorana Particles ( $v = \overline{v}$ )
- 2. demonstrate lepton number is not a symmetry of nature
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#### Signature of $0\nu\beta\beta$



#### $\beta\beta$ summed e<sup>-</sup> energy spectrum



Look for peak in the detector at the Q-value of decay.
 Good energy resolution of a detector suppresses intrinsic background from 2νββ.

#### Signature of $0\nu\beta\beta$





Look for peak in the detector at the Q-value of decay.
 Good energy resolution of a detector suppresses intrinsic background from 2νββ.

#### Search for 0v β β





$T_{1/2}^{0\nu}$	0vββ half-life	
G <sup>0</sup> v (Q,Z)	phase space factor $(\propto Q^5)$	
M0^	Nuclear Matrix Element (NME)	
m <sub>ββ</sub>	effective $\beta\beta$ neutrino mass	
m <sub>e</sub>	electron mass	

#### Search for $0\nu\beta\beta$



Decay rate:

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q,Z) |M^{0\nu}|^2$$

 ${\sf T}_{\rm I/2}{}^{\rm 0v}$  sensitivity  $\propto a\cdot\epsilon\sqrt{}$ 

$$\frac{|\langle m_{\beta\beta}\rangle|^2}{m_e^2}$$

$$\left| \frac{M \cdot t}{b \cdot \delta E} \right|$$

T <sub>1/2</sub> 0	V	0vββ half-life	
G <sup>0</sup> v (Q	$^{ m Dv}$ (Q,Z) phase space factor ( $\propto Q^5$		
M <sup>0</sup>	,	Nuclear Matrix Element (NME)	
mββ	3	effective ββ neutrino mass	
m <sub>e</sub>		electron mass	
a		isotopic abundance of source	
3		detection efficiency	
Μ		total detector mass	
b		background rate /mass/energy	
t		exposure time	
δΕ	en	ergy resolution (spectral width)	

#### Search for 0v β β

 $\langle m_{\beta\beta} \rangle$ 



Decay rate:

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q,Z) |M^{0\nu}|^2$$

 $T_{1/2}^{0\nu}$  sensitivity  $\propto a \cdot \epsilon \sqrt{\frac{M \cdot t}{b \cdot \delta E}}$ 

Source Selection/ Detector Building Strategies

- Large total mass
- Ultra-low background
- Good energy resolution
- High Q-value
- High isotopic abundanceNME

T <sub>1/2</sub> <sup>0</sup> ∨	0vββ half-life	
$G^{0\nu}(Q,Z)$	phase space factor $(\propto Q^5)$	
M0v	Nuclear Matrix Element (NME)	
mββ	effective $\beta\beta$ neutrino mass	
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a	isotopic abundance of source
3	detection efficiency
Μ	total detector mass
b	background rate /mass/energy
t	exposure time
δΕ	energy resolution (spectral width)

## Search for $0\nu\beta\beta$











KamLAND-Zen





 $T_{1/2} > 1.1 \times 10^{25}$  years Nature 510 (2014) 229–234

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#### Search for $0\nu\beta\beta$ : <sup>76</sup>Ge





#### Combined ${}^{76}GeT_{1/2} > 3.0 \times 10^{25}$ years

Phys.Rev.Lett. 110 (2013) 062502

#### Search for $0\nu\beta\beta$ : <sup>130</sup>Te



Decay rate:

$$T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q,Z) |M^{0\nu}|^2$$

 $T_{1/2}^{0\nu}$  sen

$$\frac{|\langle m_{\beta\beta}\rangle|^2}{m_e^2}$$

sitivity 
$$\propto a \cdot \epsilon \sqrt{rac{M \cdot t}{b \cdot \delta E}}$$

T <sub>1/2</sub> ⁰∨	0vββ half-life	
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M0^	Nuclear Matrix Element (NME)	
mββ	effective $\beta\beta$ neutrino mass	
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Source Selection/ **Detector Building Strategies** Large total mass Ultra-low background Good energy resolution High Q-value High isotopic abundance NME







 High isotopic abundance, low background at the Qvalue makes <sup>130</sup>Te appealing for 0vββ search.

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#### **TeO<sub>2</sub> Bolometers**



5



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Time [s]

#### Outline



- $\bigcirc$  Neutrinoless double-beta decay (0νββ) search
- CUORE : An array of TeO<sub>2</sub> bolometers
- **OUDRE-0** :  $0\nu\beta\beta$  search w/ a single CUORE tower
  - CUORE-0 : Detector
  - CUORE-0 : Resolution and Background
  - CUORE-0 : Results

## Summary

#### The CUORE 0νββ Search

CUORE-0

(20|3-20|5)



CUORE: Cryogenic Underground Observatory for Rare Events

#### Cuoricino (2003-2008)



Achieved (2015)

CUORE (2015-2020)



Achieved (2008)  $T_{1/2}^{0\nu} > 2.8 \times 10^{24} \text{ yr} (90\% \text{ C.L.})$ 

Astroparticle Physics 34 (2011) 822

#### The CUORE 0vββ Search





#### **CUORE** Collaboration





(Oct. 31, 2013 @ LNGS)





#### CUORE at LNGS





Casale San Nicola

Fonte Cerreto-(Base Funivia)

A24

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Assergi

#### The CUORE Detector





#### The CUORE Detector





#### Progress towards CUORE





#### Outline



- Solution Neutrinoless double-beta decay (0νββ) search
- CUORE : An array of TeO<sub>2</sub> bolometers
- **OVERTINE OVER SET UP A SINGLE CUORE TOWER CUORE tower** 
  - CUORE-0 : Detector
  - CUORE-0 : Resolution and Background
  - CUORE-0 : Results

#### Summary

#### CUORE-0



#### The first CUORE-like tower



#### **Detector Assembly**





Crystals are prepared & assembled into towers inside N<sub>2</sub>-fluxed glove boxes in a Class 1000 clean room.



#### **Detector Assembly**





#### **Completed Tower**











After assembly

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Attached to Cuoricino dilution refrigerator

#### The CUORE-0 Experiment





52 (13 x 4) crystals, 39 kg of TeO<sub>2</sub> (11 kg of <sup>130</sup>Te), 4 kg of copper structure.

- Validated new cleaning and assembly procedures for CUORE.
- Verified understanding on the background sources.

Eur. Phys. J. C 74, 2956 (2014)

- Tested DAQ & Analysis framework for CUORE.
- Taking 0vββ data since March 2013
   in former Cuoricino cryostat.

## **Tower Response**

Run 201388 Working Resistances

40.55

45.76

36.79

40.21

13

12

11

10

9

8

7

6

5

4

3

2

34.94

44.08

42.32

40.34

32.4

39.18



Run 201388 Working Temperatures

#### Outline



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## Summary





#### Analysis Procedure: Experimental Input







#### Calibration Spectrum





Energy resolution is evaluated for each bolometer and dataset by fitting the 2615 keV peak from <sup>208</sup>TI in the calibration data.

The obtained resolution is < 5 keV, which is the CUORE goal.</p>







- $\gamma$  background (from <sup>232</sup>Th) was not reduced since the cryostat remained the same.
- Y background (from <sup>238</sup>U chain) was reduced by a factor of 2.5 due to better radon control.
- $\alpha$  background from copper surface and crystal surface was reduced by a factor of Background Paper 6.5 thanks to the new detector surface treatment. in Preparation!
- Demonstrate CUORE sensitivity goal is within reach.

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	Background rate [counts/keV/kg/y]		signal eff. [%]
	0vββ region	α region (excl. peak)	(detector+cuts)
Cuoricino	0.169 ± 0.006	0.110 ± 0.001	82.8±1.1
CUORE-0	0.058 ± 0.011	0.016 ± 0.001	81.3±0.6

#### Outline



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#### Summary

## Blinding 0v BB Region

dN/dE

Region of Interest was blinded by "salting": A small (and *blinded*) fraction of the events within  $\pm 10$  keV in <sup>208</sup>Tl photopeak are exchanged with events within  $\pm 10$  keV of the  $0\nu\beta\beta$  Q-value to produce a *fake* peak.



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Background at ROI can be characterized without biasing 0vββ analysis.









#### Fit to the Unblinded ROI



Simultaneous unbinned extended ML fit to range [2470,2570] keV

- Fit function has 3 components:
  - I. Calibration-derived lineshape modeling posited fixed at 2527.5 keV
  - 2. Calibration-derived lineshape modeling Co peak floated around 2505 keV
  - 3. Continuum background



#### Fit to the Unblinded ROI



Fitted background:  $0.058 \pm 0.004 \text{ (stat.)} \pm 0.002 \text{ (syst.)} \text{ counts/keV/kg/yr}$ Best-fit decay rate:  $\Gamma^{0\nu\beta\beta}(^{130}\text{Te}) = 0.01 \pm 0.12 \text{ (stat.)} \pm 0.01 \text{ (syst.)} \times 10^{-24} \text{ yr}^{-1}$ 

![](_page_49_Picture_0.jpeg)

#### Fit to the Unblinded ROI

![](_page_49_Figure_2.jpeg)

 $\Gamma^{0\nu\beta\beta}(^{130}\text{Te}) < 0.25 \times 10^{-24} \text{ yr}^{-1} (90\% \text{ C.L., statistics only})$  $T_{1/2}^{0\nu\beta\beta}(^{130}\text{Te}) > 2.7 \times 10^{24} \text{ yr} (90\% \text{ C.L., statistics only})$ 

#### **Systematics**

![](_page_50_Picture_1.jpeg)

TABLE I. Systematic uncertainties on  $\Gamma_{0\nu}$  in the limit of zero signal (Additive) and as a percentage of nonzero signal (Scaling).

	Additive $(10^{-24})$	$y^{-1}$ ) Scaling (%)
Lineshape	0.007	1.3
Energy resolution	0.006	2.3
Fit bias	0.006	0.15
Energy scale	0.005	0.4
Bkg function	0.004	0.8
Selection efficiency	0.	7%

- For each systematic, we run toy MC exps. to evaluate bias on fitted  $0\nu\beta\beta$  rate.
- Bias is parameterized as p0 + p1xΓ, where p0 = "additive" and p1="scaling"
- Signal lineshape: Used variety of different line shapes to model signal
- Energy resolution: Apply 1.05 ± 0.05 correction to calibration-derived resolution
- Fit bias: Effect of using unbanned extended ML fit to extract values
- Energy scale: Assign 0.12 keV uncertainty derived from peak residuals in physics spectrum
- **Bkg function**: Choices of 0-, 1-, 2- order polynomial.

#### **Systematics**

![](_page_51_Picture_1.jpeg)

TABLE I. Systematic uncertainties on  $\Gamma_{0\nu}$  in the limit of zero signal (Additive) and as a percentage of nonzero signal (Scaling).

	Additive $(10^{-24})$	$y^{-1}$ ) Scaling (%)
Lineshape	0.007	1.3
Energy resolution	0.006	2.3
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Bkg function	0.004	0.8
Selection efficiency	0.	7%

We find **no evidence** for  $0\nu\beta\beta$  of <sup>130</sup>Te (report the Bayesian limits)

 $\Gamma^{0\nu\beta\beta}$  (<sup>130</sup>Te) < 0.25 × 10<sup>-24</sup> yr<sup>-1</sup> (90% C.L., stat.+sys.)

 $T_{1/2}^{0\nu\beta\beta}(^{130}\text{Te}) > 2.7 \times 10^{24} \text{ yr} (90\% \text{ C.L., stat.+sys.})$ 

## Combining Cuoricino & CUORE+0

![](_page_52_Figure_1.jpeg)

Combining the CUORE-0 result with the Cuoricino result from 19.75 kg-yr of <sup>130</sup>Te exposure yields the Bayesian lower limit:

 $T_{1/2}^{0\nu\beta\beta}(^{130}\text{Te}) > 4.0 \times 10^{24} \text{ yr} (90\% \text{ C.L., stat.+sys.})$ 

arXiv:1504.02454 Submitted to PRL

![](_page_53_Figure_0.jpeg)

![](_page_53_Figure_1.jpeg)

![](_page_53_Figure_2.jpeg)

Including additional Shell-Model NME

 $\langle \mathbf{m}_{\beta\beta} \rangle < 270 - 760 \text{ meV}$ 

1) IBM-2 (PRC 91, 034304 (2015))
 2) QRPA (PRC 87, 045501 (2013))
 3) pnQRPA (PRC 024613 (2015))
 4) Shell Model (PRC 91, 024309 (2015))
 5) ISM (NPA 818, 139 (2009))
 6) EDF (PRL 105, 252503 (2010))

#### CUORE: Background Budget

![](_page_54_Picture_1.jpeg)

![](_page_54_Figure_2.jpeg)

#### **CUORE:** Sensitivity

![](_page_55_Picture_1.jpeg)

![](_page_55_Figure_2.jpeg)

- Assumptions: 5 keV FWHM ROI resolution ( $\delta E$ ), background rate (b) of 0.01 counts/(keV kg yr) arXiv:1109.0494
- **5** years of live time.

#### Summary

![](_page_56_Picture_1.jpeg)

- Observation of  $0\nu\beta\beta$  will establish that neutrinos are Majorana particles.
- TeO<sub>2</sub> bolometers offer a well-established and competitive technique to search for  $0\nu\beta\beta$ .
- CUORE-0 and Cuoricino, the experiments on the way to CUORE, did not find evidence of  $0\nu\beta\beta$  of <sup>130</sup>Te.
- CUORE-0, the first CUORE-like tower currently operating at LNGS, demonstrated background suppression and resolution improvements, i.e., achieved goals for CUORE.
- CUORE, the largest cryogenic detector using TeO<sub>2</sub> bolometers with 206 kg of <sup>130</sup>Te mass, completed detector construction and commissioning of the cryogenic system along with infrastructure is well underway.
- CUORE is scheduled to start data-taking in late 2015 and various R&D projects are on-going for searches beyond CUORE.